

Theo.4. Statistical Physics

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Lecture 21.

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January 21, 2024.

Last time: Two-level system & ideal gas in canonical ensemble

Single vs. multi particle partition fun.

Today: Gibbs canonical ensemble

Grand canonical ensemble.

From last time, we discussed the canonical ensemble, where a system is thermally coupled to a reservoir at constant T . By analyzing the Helmholtz free energy via the partition fun. expression

$$F(T, \vec{x}) = -k_B T \ln Z(T, \vec{x}),$$

we can calculate (any) desired thermodynamic property of the system, such as avg. energy, heat capacity, pressure, chemical potential, etc.

Now, we construct a new Stat. mech system, where a system is allowed to change internal energy via heat and work. Macrostate $M \equiv (T, \vec{J})$ specifies temperature and forces. System maintained at constant force via external elements (pistons or magnets). Energy of the combined system is $H - \vec{J} \cdot \vec{x}$ including work done against the forces. Work done on the system is $+\vec{J} \cdot \vec{x}$.

Microstates of combined system has canonical prob.

$$p(\mu_s, \vec{x}) = \exp[-\beta H(\mu_s) + \beta \vec{J} \cdot \vec{x}] / \mathcal{Z}(T, N, \vec{J})$$

with Gibbs partition fun $\mathcal{Z}(N, T, \vec{J}) = \sum_{\mu_s, \vec{x}} \exp[\beta \vec{J} \cdot \vec{x} - \beta H(\mu_s)]$

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Note RHS has no N dependence: there is no chemical potential (will be considered for grand canonical ensemble ~~potential~~).

It is easy to calculate the equilibrium exp. values of the displacements $\vec{x}_i$:

$$\langle \vec{x}_i \rangle = k_B T \frac{\partial \ln \tilde{Z}}{\partial J_i}$$

Since $\vec{x} = -\frac{\partial G}{\partial \vec{J}} \Rightarrow$ identify $G(N, T, \vec{J}) = -k_B T \ln \tilde{Z}$,

where $G \equiv E - TS - \vec{x} \cdot \vec{J}$. Can also derive by maximizing probability wrt. $\vec{x}$.

Enthalpy $H \equiv E - \vec{x} \cdot \vec{J}$ obtained by

$$-\frac{\partial \ln \tilde{Z}}{\partial \beta} = \langle H - \vec{x} \cdot \vec{J} \rangle = H.$$

Heat capacities at const. force (including work done against ext forces) arise from enthalpy as $C_J = \frac{\partial H}{\partial T}$.

Example Ideal gas in isobaric ensemble.

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$M \equiv (N, T, P)$ macrostate.

Microstate $\mu \equiv \{ \vec{p}_i, \vec{q}_i \}$ with volume V .

$$P(\{ \vec{p}_i, \vec{q}_i \}, V) = \frac{1}{\tilde{Z}} \exp \left[-\beta \sum_{i=1}^N \frac{p_i^2}{2m} - \beta P V \right]$$

$\begin{cases} 1 & \text{for } \{q_i\} \text{ in box } V \\ 0 & \text{otherwise.} \end{cases}$

Note) again, not uniform prob.

distribution, but instead weighted by Boltzmann factor $\exp(-\frac{E}{k_B T})$ and $\exp(-\frac{PV}{k_B T})$.

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Normalization factor: as usual, total # of ~~microstates~~ probabilities.

$$\begin{aligned}
\mathcal{Z}(N, T, P) &= \int_0^\infty dV e^{-\beta PV} \int \frac{1}{N!} \prod_{i=1}^N \frac{d^3 \vec{p}_i d^3 \vec{q}_i}{h^3} \exp\left[-\beta \sum_{i=1}^N \frac{p_i^2}{2m}\right] \\
&= \int_0^\infty dV V^N e^{-\beta PV} \frac{1}{N! \lambda(T)^{3N}} \\
&= \frac{1}{(\beta P)^{N+1} \lambda(T)^{3N}}
\end{aligned}$$

where $\lambda(T) = \frac{h}{\sqrt{2\pi m k_B T}}$ is the thermal wavelength from before.

Ignore non-extensive contributions

$$\begin{aligned}
G &= -k_B T \ln \mathcal{Z} \\
&\approx N k_B T \left[\ln P - \frac{5}{2} \ln(k_B T) + \frac{3}{2} \ln\left(\frac{h^2}{2\pi m}\right) \right]
\end{aligned}$$

Given $dG = -SdT + VdP + \mu dN$,

volume is $V = \left. \frac{\partial G}{\partial P} \right|_{T, N} = \frac{N k_B T}{P} \Rightarrow PV = N k_B T$.

Enthalpy is $H = \langle E + PV \rangle$

$$H = - \frac{\partial \ln \mathcal{Z}}{\partial \beta} = \frac{5}{2} N k_B T,$$

$$C_p = \frac{dH}{dT} = \frac{5}{2} N k_B.$$

Example 2.

Spins in ext. magnetic field $\vec{B}$.

Work done against magnetic field + internal Hamiltonian gives Gibbs partition fun.

$$\mathcal{Z}(N, T, B) = \text{tr} \left[\exp(-\beta \mathcal{H} + \beta \vec{B} \cdot \vec{M}) \right]$$

↑
"sum over all dof"

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Here, the d.o.f. can be taken as spin $\frac{1}{2}$ particles with two possible projections of spin along magnetic field. Microstate of N spins is a set of Ising $\{\sigma_i = \pm 1\}$ variables.

Corresponding magnetization along field direction

$$M = \mu_0 \sum_{i=1}^N \sigma_i, \quad \mu_0 \text{ is microscopic mag. moment.}$$

If $H=0$ (no interactions among spins),

$$P(\{\sigma_i\}) = \frac{1}{\mathcal{Z}} \left[\beta \mu_0 B \sum_{i=1}^N \sigma_i \right]$$

(cf. two-level system from earlier)

Gibbs partition fun.

Recall: $\cosh(x) = \frac{1}{2}(e^x + e^{-x})$

$$\mathcal{Z}(N, T, B) = [2 \cosh(\beta \mu_0 B)]^N$$

Gibbs free energy $G = -k_B T \ln \mathcal{Z} = -N k_B T \ln [2 \cosh(\beta \mu_0 B)]$

Avg. magnetization

Recall:

$$M = -\frac{\partial G}{\partial B} = N \mu_0 \tanh(\beta \mu_0 B)$$

$$\frac{d}{dx} \cosh x = \sinh x$$

Can expand for small B & obtain well-known Curie Law of magnetic susceptibility

$$\chi(T) = \left. \frac{\partial M}{\partial B} \right|_{B=0} = \frac{N \mu_0^2}{k_B T}$$

Enthalpy $H = \langle H - BM \rangle$
 $= -BM$

$$\downarrow C_B = -B \frac{\partial M}{\partial T}$$

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The final ensemble to consider is the grand ~~can~~ canonical ensemble, which allows chemical work but not mechanical work. Hence, system macrostate fixed by chemical potential μ , $M \equiv (T, \mu, \vec{x})$. Microstates μ_s contain indefinite # of particles.

$$p(\mu_s) = \frac{\exp[\beta \mu N(\mu_s) - \beta H(\mu_s)]}{Q}$$

Grand partition fcn. $Q(T, \mu, \vec{x}) = \sum_{\mu_s} \exp[\beta \mu N(\mu_s) - \beta H(\mu_s)]$

Easier to analyze by separating a sum over # of given particles,

$$Q(T, \mu, \vec{x}) = \sum_{N=0}^{\infty} e^{\beta \mu N} \sum_{(\mu_s | N)} e^{-\beta H_N(\mu_s)}$$

Can recognize that restricted sums $\sum_{(\mu_s | N)}$ are regular N -particle partition fcn's. Each term in Q is total weight of all microstates of N particles.

Unconditional prob. of finding N particles is

$$p(N) = \frac{e^{\beta \mu N} Z(T, N, \vec{x})}{Q(T, \mu, \vec{x})}$$

$$\langle N \rangle = \frac{1}{Q} \frac{\partial}{\partial (\beta \mu)} Q = \frac{\partial \ln Q}{\partial (\beta \mu)}$$

Can check number fluctuations are related to variance

$$\langle N^2 \rangle_c = \frac{\partial \langle N \rangle}{\partial (\beta \mu)} \Rightarrow \text{Relative number fluct. vanish in thermodynamic limit.}$$

Sharpness for $N \Rightarrow$ approx. sum by largest term
at $N = N^* \simeq \langle N \rangle$.

$$\begin{aligned} Q(T, \mu, \vec{x}) &= \lim_{N \rightarrow \infty} \sum_{N=0}^{\infty} e^{\beta \mu N} Z(T, N, \vec{x}) \\ &= e^{\beta \mu N^*} Z(T, N^*, \vec{x}) \\ &= e^{\beta \mu N^* - \beta F} = e^{-\beta G} \end{aligned}$$

where $G(T, \mu, \vec{x}) = E - TS - \mu N = F - \mu N$
 $= -k_B T \ln Q$

is grand potential.

We know $dG = -SdT - Nd\mu + \vec{J} \cdot d\vec{x}$.

Again, example in Kardar for ideal gas, p.119-120.

Get $Q(T, \mu, V) = \exp \left[\beta \mu \frac{V}{\lambda^3} \right],$

$$G(T, \mu, V) = -k_B T \left(e^{\beta \mu} \frac{V}{\lambda^3} \right)$$

$$\mu = k_B T \ln \left(\frac{p \lambda^3}{k_B T} \right)$$

Next time: Legendre transforms.

Comments about interacting systems.