

October 18, 2019.

Metric is always
(+, -, -, -)

Introduction + Motivation

Distinction between studying "real" phenomena + "aesthetic"/ theory problems.

Real = DM, baryogenesis, quantum gravity

Aesthetic = "Fine-tuning" hierarchy problem, Θ -QCD, cosmological constant

Similar to distinction between bottom-up + top-down approaches to model-building.

Bottom-up: Add particles for ν -mass, DM, flavor puzzles in $g-2$ or R_K, K^* anomalies.

Parametrize NP effects by EFT, agnostic about possible coupling patterns in UV physics.

Top-down: Impose new symmetry; spurion analysis can be very useful.

Will adopt top-down perspective: Add as new symmetry to $\mathcal{L}$.

Should already know how to analyze global symmetries + gauge symmetries.

Recall that these symmetries (if non-anomalous), leave action invariant, leading to conserved Noether currents + charges commute with Hamiltonian, ~~and~~ making the charges good quantum numbers. ~~the factors the~~
~~the factors the~~

Given that you can readily build more complicated

QFTs from adding more + more gauge + global symmetries with more dofs, what about geometric extensions?

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Coleman-Mandula theorem:

Phys. Rev. 159,

Shifman,
(chap. 46,
p. 413.

Following Witten; (Intro to SUSY,

ed. Zichichi, The Unity of the
Fundamental Interactions, 1983)

1251 (1967)

For theories with nontrivial S matrix, no geometric extensions of Poincaré algebra are possible.

Can only have have 4 P_μ (energy-momentum operator) and six Lorentz transformations $M_{\mu\nu}$. Rest must be scalars such as EM charge.

~~Wess & Bagger~~

Wess &
Bagger,
p. 4

Explicitly, Coleman-Mandula assumes:

- 1.) S-matrix is based on local, relativistic QFT in 4D
- 2.) There are ^{only} finite # of diff. particles associated with one-particle states of a given mass.
- 3.) There is an energy gap between the vacuum & one particle states.

Then the most general Lie algebra of symmetries of S-matrix contains ^{energy-momentum op.} P_μ , Lorentz rotation generator $M_{\mu\nu}$, & finite # of Lorentz scalar operators B_i . Further, B_i must belong to Lie algebra of a compact Lie group.

Gelfand &
Lichtman
(1970)

Loophole: Generalize Lie algebra to graded Lie algebra, which allows anticommutators & commutators. Can then introduce spinorial generators with anticommutation relations.

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Recall Lorentz boosts, rotations + translations form $SU(2, \mathbb{C})$

Poincaré group.

Lorentz group is $SO(1, 3)$, Poincaré isPeskin,
3.1

$$\vec{J} = \vec{x} \times \vec{p} = \vec{x} \times (-i\vec{\nabla})$$

$$\mathbb{R}^{1,3} \times SO(1, 3)$$

generalize to $J^{\mu\nu} = i(x^\mu \partial^\nu - x^\nu \partial^\mu)$, antisymmetric tensor

(commutator is

$$[J^{\mu\nu}, J^{\rho\sigma}] = i(g^{\mu\rho} J^{\nu\sigma} - g^{\mu\sigma} J^{\nu\rho} - g^{\nu\rho} J^{\mu\sigma} + g^{\nu\sigma} J^{\mu\rho})$$

(can define generators of rotations + boosts as

$$L_i = \frac{1}{2} \epsilon^{ijk} J^{jk}$$

$$K_i = J^{0i} \quad \text{for } i, j, k = 1, 2, 3.$$

Infinitesimal Lorentz transformation is

$$\Phi \rightarrow (1 - i\vec{\theta} \cdot \vec{L} - i\vec{\beta} \cdot \vec{K}) \Phi$$

Finite dim. reps are those with integer or half-integer values for angular momentum.

Poincaré adds momentum generators to Lorentz.

$$[p^\mu, p^\nu] = 0$$

$$[J^{\mu\nu}, p^\rho] = i(-g^{\mu\rho} p^\nu + g^{\nu\rho} p^\mu)$$

Correspond to conservation of energy + momentum + angular momentum.

Shifman
47.2(In Weyl representation) Introduce 4 new fermionic ~~conser~~ supercharges, $Q_\alpha, \bar{Q}^{\dot{\alpha}}$ where $\alpha, \dot{\alpha} = 1, 2$.

$$[p_\mu, Q_\alpha] = [p_\mu, \bar{Q}^{\dot{\alpha}}] = 0$$

$$[J^{\mu\nu}, Q_\alpha] = i(\sigma^{\mu\nu})_\alpha{}^\beta Q_\beta$$

$$[J^{\mu\nu}, \bar{Q}^{\dot{\alpha}}] = i(\bar{\sigma}^{\mu\nu})^{\dot{\alpha}}{}_{\dot{\beta}} \bar{Q}^{\dot{\beta}}$$

Since $Q_\alpha, \bar{Q}^{\dot{\alpha}}$ transform as Weyl spinors:

$$\begin{aligned} \sigma^{\mu\nu} &\equiv \frac{1}{4}(\sigma^\mu \bar{\sigma}^\nu - \sigma^\nu \bar{\sigma}^\mu) = \left(-\frac{1}{2}\vec{\sigma}, \frac{1}{2}\vec{\sigma} \right) \\ \bar{\sigma}^{\mu\nu} &\equiv \frac{1}{4}(\bar{\sigma}^\mu \sigma^\nu - \bar{\sigma}^\nu \sigma^\mu) = \left(\frac{1}{2}\vec{\sigma}, \frac{1}{2}\vec{\sigma} \right) \end{aligned} \quad \left. \begin{aligned} F^{\mu\nu} &= (-\vec{E}, \vec{B}) \\ F_{\mu\nu} &= (\vec{E}, \vec{B}) \end{aligned} \right\}$$

Key point: Need anticommutators $\{Q_\alpha, \bar{Q}_\beta\}$ and $\{Q_\alpha, Q_\beta\}$ to close the algebra.

$$\{Q_\alpha, \bar{Q}_\beta\} = 2 p_\mu (\sigma^\mu)_{\alpha\beta} = 2 p_{\alpha\beta}$$

And choose $\{Q_\alpha, Q_\beta\} = \{\bar{Q}_\alpha, \bar{Q}_\beta\} = 0$

Alternative is $\{Q_\alpha, Q_\beta\} = C_{\alpha\beta}$, where $C_{\alpha\beta}$ is a triplet of central charges which commute w/ all other generators, like topological currents + charges.

Comments:

1.) Can also have $N=2, N=4$ SUSY in 4D. Put $\{Q_\alpha^A, \bar{Q}_\beta^B\} = 2 \sigma_{\alpha\beta}^\mu p_\mu \delta^A_B$ for $A, B = 1, \dots, N$.

2.) Coleman + Mandula theorem briefly:

Only symmetry of S-matrix that includes Poincaré is product of Poincaré and internal symmetry group.

Suppose false: Additional symmetry generators

transform as Lorentz tensors would ~~be~~ over-constrain

S-matrix. Recall in 2-2 scattering, S-matrix is fcn. only of one variable, scattering angle.

If tensor current is conserved, then angle is no ~~longer~~ longer continuous + S-matrix is not ~~at~~ analytic.

SUSY evades Coleman + Mandula with anticommutators + spinor generators.

Terning,
p. 5

3.) SUSY algebra is invariant under multiplication of Q_α by a phase, so there is one linear combination of $U(1)$ charges, "R-charge" that does not commute with $Q + Q^\dagger$:

$$[Q_\alpha, R] = Q_\alpha, \quad [Q^\dagger_\alpha, R] = -Q^\dagger_\alpha.$$

This is R-symmetry (later, gives origin of R-parity.)

The SUSY algebra extends the Poincaré algebra to a super-Poincaré algebra. Fundamentally distinct from other kinds of symmetries that you can consider, like gauge + global symmetries.

Witten p. 422 In non-SUSY field theories, we had fields of spin 0, $1/2$, $\neq 1$ that depend ^{locally} on spacetime x^μ .

Just as ~~the~~ Energy-momentum tensor generates translations in 4D spacetime, anticommuting supercharges should generate "super" translations in an anticommuting space.

Introduce $\theta^\alpha, \bar{\theta}^{\dot{\alpha}}$ as four anti-commuting "fermionic" dimensions of superspace. Recall $\alpha, \dot{\alpha} = 1, 2$.

Group element of $N=1$ superalgebra is then

$$G(x^\mu, \theta, \bar{\theta}) = \exp(i(\theta Q + \bar{\theta} \bar{Q} - x^\mu p_\mu))$$

with Grassmann variables $\theta^\alpha, \bar{\theta}^{\dot{\alpha}} \equiv (\theta^\alpha)^*$

Recall Grassmann variables (from fermionic path integral approach to QFT) anti commute:

$$\{\theta^\alpha, \theta^\beta\} = \{\bar{\theta}^{\dot{\alpha}}, \bar{\theta}^{\dot{\beta}}\} = \{\theta^\alpha, \bar{\theta}^{\dot{\beta}}\} = 0$$

$$\left\{ \frac{\partial}{\partial \theta^\alpha}, \frac{\partial}{\partial \theta^\beta} \right\} = \left\{ \frac{\partial}{\partial \bar{\theta}^{\dot{\alpha}}}, \frac{\partial}{\partial \bar{\theta}^{\dot{\beta}}} \right\} = \left\{ \frac{\partial}{\partial \theta^\alpha}, \frac{\partial}{\partial \bar{\theta}^{\dot{\beta}}} \right\} = 0$$

Peskin,
p. 302

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What are the reps. in superspace?

Will see that under SUSY transformations, scalars transform into fermions & vice-versa.

(Consider how the group elements in $G(x^\mu, \theta, \bar{\theta})$ act on $G(a^\mu, \epsilon, \bar{\epsilon})$.

$$G(x^\mu, \theta, \bar{\theta}) G(a^\mu, \epsilon, \bar{\epsilon})$$

Exercise: $= G(x^\mu + a^\mu + i\epsilon\sigma^\mu\bar{\theta} - i\theta\sigma^\mu\bar{\epsilon}, \theta + \epsilon, \bar{\theta} + \bar{\epsilon})$

which follows from Hausdorff formula

~~Exercise~~ $e^A e^B = \exp(A + B + \frac{1}{2}[A, B] + \dots)$

We identify $x^\mu \rightarrow x^\mu + \delta x^\mu$

$$\text{with } \delta x^\mu \equiv \delta\left(\frac{1}{2}x_{\alpha\dot{\beta}}(\bar{\sigma}^\mu)^{\dot{\beta}\alpha}\right) \\ = \frac{1}{2}(\delta x_{\alpha\dot{\beta}})(\bar{\sigma}^\mu)^{\dot{\beta}\alpha}$$

$$\text{with } \delta x_{\alpha\dot{\beta}} = -2i\theta_\alpha\bar{\epsilon}_{\dot{\beta}} - 2i\bar{\theta}_{\dot{\beta}}\epsilon_\alpha$$

$$\text{and } \delta\theta^\alpha = \epsilon^\alpha, \quad \delta\bar{\theta}^{\dot{\alpha}} = \bar{\epsilon}^{\dot{\alpha}}$$

Aside: Weyl spinors and 2-component spinor notation

$$\Psi = \begin{pmatrix} \xi \\ \eta \end{pmatrix}$$

$$P_L \Psi = \xi \quad \text{LH-spinor}$$

$$P_R \Psi = \eta \quad \text{RH-spinor}$$

Peskin,
p. 44.

$$\xi_L \rightarrow \exp\left(1 - i\vec{\alpha} \cdot \frac{\vec{\sigma}}{2} - \vec{\beta} \cdot \frac{\vec{\sigma}}{2}\right) \xi_L$$

$$\eta_R \rightarrow \exp\left(1 - i\vec{\alpha} \cdot \frac{\vec{\sigma}}{2} + \vec{\beta} \cdot \frac{\vec{\sigma}}{2}\right) \eta_R$$

Infinitesimal rotations $\vec{\alpha}$, boosts $\vec{\beta}$.

Recall $\sigma^2 \vec{\sigma}^* = -\vec{\sigma} \sigma^2$ so $\sigma^2 (\xi_L^*)$ transforms as a RH-spinor.

Notation: α, β are spinor indices for LH Weyl spinors.
 $\dot{\alpha}, \dot{\beta}$ are spinor indices for RH Weyl spinors.

To keep track of upper + lower indices, use 7

$$\epsilon^{\alpha\beta} = -\epsilon^{\beta\alpha}, \epsilon^{12} = 1 = -\epsilon_{12}$$

$$\epsilon^{\dot{\alpha}\dot{\beta}} = -\epsilon^{\dot{\beta}\dot{\alpha}}, \epsilon^{\dot{1}\dot{2}} = 1 = -\epsilon_{\dot{1}\dot{2}}$$

Contracting spinors:

$$\eta\chi \equiv \eta^\alpha \chi_\alpha, \quad \bar{\eta}\bar{\chi} \equiv \bar{\eta}_{\dot{\alpha}} \bar{\chi}^{\dot{\alpha}}$$

where ~~$\chi_\alpha = (\chi^\alpha)^*$~~

$$\bar{\chi}_{\dot{\alpha}} \equiv (\chi_\alpha)^*, \quad \bar{\eta}^{\dot{\alpha}} \equiv (\eta^\alpha)^*$$

Can show

$$\begin{aligned} (\eta\chi)^\dagger &= (\eta^\alpha \chi_\alpha)^\dagger \\ &= (\chi_\alpha)^* (\eta^\alpha)^* \\ &= \bar{\chi} \bar{\eta} \end{aligned}$$

We also have fermion bilinears

$$\chi^\alpha \chi_\beta = -\frac{1}{2} \epsilon^{\alpha\beta} \chi^2, \quad \chi_\alpha \chi_\beta = +\frac{1}{2} \epsilon_{\alpha\beta} \chi^2$$

$$\bar{\chi}^{\dot{\alpha}} \bar{\chi}^{\dot{\beta}} = \frac{1}{2} \epsilon^{\dot{\alpha}\dot{\beta}} \bar{\chi}^2, \quad \bar{\chi}_{\dot{\alpha}} \bar{\chi}_{\dot{\beta}} = -\frac{1}{2} \epsilon_{\dot{\alpha}\dot{\beta}} \bar{\chi}^2$$

To extend to vectors, ~~use~~ $(\frac{1}{2}, \frac{1}{2})$ rep. of Lorentz,
use $(\sigma^\mu)_{\alpha\dot{\beta}} = \{1, \vec{\sigma}\}_{\alpha\dot{\beta}}$ for convolution.

$$A_\mu \equiv (\Lambda^0, -\vec{\Lambda})$$

$$\Lambda^\mu = \frac{1}{2} \Lambda_{\alpha\dot{\beta}} (\bar{\sigma}^\mu)^{\dot{\alpha}\alpha}, \quad \Lambda_{\alpha\dot{\beta}} = A_\mu (\sigma^\mu)_{\alpha\dot{\beta}}$$

Consider $(\sigma^\mu)_{\alpha\dot{\beta}}$ RH + $(\bar{\sigma}^\mu)^{\dot{\alpha}\alpha}$ LH.

Recall Dirac equation:

$$\bar{\Psi} i \not{\partial} \Psi - m \bar{\Psi} \Psi = 0$$

$$\gamma^\mu = \begin{pmatrix} 0 & \sigma^\mu \\ \bar{\sigma}^\mu & 0 \end{pmatrix}$$

For free, $\not{\partial} \Rightarrow \not{\partial} \Rightarrow$

$$\begin{pmatrix} -m & i\sigma \cdot \partial \\ i\bar{\sigma} \cdot \partial & -m \end{pmatrix} \begin{pmatrix} \psi_L \\ \psi_R \end{pmatrix} = 0$$

$$\text{or } i\sigma \cdot \partial \psi_R = m \psi_L$$

$$i\bar{\sigma} \cdot \partial \psi_L = m \psi_R$$

⑧

(can extend this distinction to superspace.

Two invariant subspaces:

$$\{x_L^m, \theta^\alpha\}, \quad \delta\theta^\alpha = \epsilon^\alpha, \quad \delta(x_L)_{\alpha\dot{\alpha}} = -4i\theta_\alpha \bar{\epsilon}_{\dot{\alpha}}$$

$$\{x_R^m, \bar{\theta}^{\dot{\alpha}}\}, \quad \delta\bar{\theta}^{\dot{\alpha}} = \bar{\epsilon}^{\dot{\alpha}}, \quad \delta(x_R)_{\alpha\dot{\alpha}} = -4i\bar{\theta}_{\dot{\alpha}} \epsilon_\alpha$$

$$\text{with } (x_L)_{\alpha\dot{\alpha}} = x_{\alpha\dot{\alpha}} - 2i\theta_\alpha \bar{\theta}_{\dot{\alpha}}$$

$$(x_R)_{\alpha\dot{\alpha}} = x_{\alpha\dot{\alpha}} + 2i\theta_\alpha \bar{\theta}_{\dot{\alpha}}$$

Will also use vectorial notation

$$x_L^m = x^m - i\theta^\alpha (\sigma^m)_{\alpha\dot{\alpha}} \bar{\theta}^{\dot{\alpha}},$$

$$x_R^m = x^m + i\theta^\alpha (\sigma^m)_{\alpha\dot{\alpha}} \bar{\theta}^{\dot{\alpha}}$$

In SUSY, we have superfields as ~~scalar~~ fns. on superspace coordinates. Expanding superfield in powers of supervariables $\theta^\alpha + \bar{\theta}^{\dot{\alpha}}$, we get the regular fields as ~~scalar~~ components of the superfield.

Remark: Expansion is finite since Taylor expansion over Grassmann variables is finite. Highest term is $\theta^2 \bar{\theta}^2 \equiv \theta^\alpha \theta_\alpha \bar{\theta}_{\dot{\alpha}} \bar{\theta}^{\dot{\alpha}}$

Most general superfield is

$$S(x, \theta, \bar{\theta}) = \phi + \theta\psi + \bar{\theta}\bar{\chi} + \theta^2 F + \bar{\theta}^2 G$$

$$+ \theta^\alpha A_{\alpha\dot{\beta}} \bar{\theta}^{\dot{\beta}} + \theta^2 (\bar{\theta}\mathcal{T})$$

$$+ \bar{\theta}^2 (\theta\rho) + \bar{\theta}^2 \bar{\theta}^2 \mathcal{D}, \text{ with}$$

$\phi, \psi, \bar{\chi}, \dots, \mathcal{D}$ as component fields depending

only on x^m .

Next time: ~~Reps. of SUSY algebra~~

Reps. of SUSY algebra based on $S(x, \theta, \bar{\theta})$.