

Lecture 17

Quantum Field Theory II: The SM + EW Theory

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Last time: EFTs as top-down + bottom-up tool; SM EFT

Today: EFTs in everyday use

DM EFTs & simplified models
ν-physics + seesaw physics

Recall EFTs are defined by low-energy set of dynamical fields, gauge + global symmetries, + truncation order of higher dimensional operators.

If we consider Dark Matter, then we need to introduce a new deg. of freedom to the SM that can serve as a DM candidate. The main requirements for particle dark matter are:

① Cosmologically stable: either does not decay or lifetime is longer than the age of the universe.

② Electrically neutral: Given DM in the galactic halo, an electric charge for DM would give photon interactions in the CMB that are excluded. Exception: microcharges.

③ Color neutral: Similar to #2, we expect a local density of DM in our solar system. Then, surface-based or underground detectors would expect to see nuclear recoils from new colored states.

(2)

④ Relic density: $\Omega h^2 \sim 0.12$ The energy budget of the universe (from CMB measurements) has

$$\text{been fit to } \Omega_m = 0.311 \pm 0.006$$

$$\Omega_\Lambda = 0.687 \pm 0.006$$

[PDG,

Cosmological Params.,

Table 25.1, 2020]

Ω_m includes baryonic & dark matter

baryonic matter $\Omega_b h^2 = 0.02242 \pm 0.00014$

$$h = 0.677 \pm 0.004$$

cold DM $\Omega_c h^2 = 0.1193 \pm 0.0009$

This is in the Λ CDM standard cosmological model, which is beyond the scope of these lectures.

⑤ Mass is 10^{-20} eV to $100 M_\odot$

Thermal DM is $\sim$ MeV to 10^{59} GeV

$$M_\odot = 1.989 \cdot 10^{33} \text{ g}$$

$$m_p = 1.67 \cdot 10^{-24} \text{ g}$$

10^{-20} eV $\sim$ fuzzy DM

$$M_\odot \sim 10^{57} m_p \sim 10^{57} \text{ GeV}$$

10^{59} eV $\sim$ black holes (primordial)

To construct an EFT, we now also assume

① DM is an $SU(2)$ singlet.

② DM is a Dirac fermion (can also be scalar or Majorana ferm)

③ DM is odd under $\mathbb{Z}_2$ sym, SM is even $\Rightarrow$ ensures DM stability
Write possible $\mathcal{L}$ interactions up to dim. 6:

$$\mathcal{L}_{\text{tot}} = \mathcal{L}_{\text{SM}} + \mathcal{L}_{\text{DM}} + \mathcal{L}_{\text{DM+SM EFT}}$$

$$\mathcal{L}_{\text{DM}} = \bar{\chi} i \not{\partial} \chi - m \bar{\chi} \chi$$

$$\mathcal{L}_{\text{DM-SM EFT}} = \left(\begin{array}{c} \text{SM gauge} \\ \text{invariant,} \\ \text{Lorentz} \\ \text{covariant} \end{array} \right) \otimes \left(\begin{array}{c} \text{DM gauge} \\ \text{invariant,} \\ \text{Lorentz} \\ \text{covariant} \end{array} \right)$$

(3)

Note that fermion bilinears mixing SM + DM contractions can be generated by Fierz identities.

For simplicity, work in SM broken phase:

So we have only $SU(3)_c \times U(1)_{em}$ gauge symmetries. Starting with the DM side, we have the usual 5 fermion bilinears:

$$\bar{\chi}\chi, \quad \bar{\chi}\gamma^5\chi, \quad \bar{\chi}\gamma^\mu\chi, \quad \bar{\chi}\gamma^\mu\gamma^5\chi, \quad \bar{\chi}\sigma^{\mu\nu}\chi.$$

These can be matched to corresponding quark bilinears, subject to overall Lorentz invariance.

$$\begin{aligned} \mathcal{L}_{DM-SM}^{EFT} = & \frac{1}{\Lambda^2} \bar{q}q \bar{\chi}\chi + \frac{1}{\Lambda^2} \bar{q}\gamma^5 q \bar{\chi}\chi + \frac{1}{\Lambda^2} \bar{q}q \bar{\chi}\gamma^5\chi \\ & + \frac{1}{\Lambda^2} \bar{q}\gamma^5 q \bar{\chi}\gamma^5\chi \\ & + \frac{1}{\Lambda^2} \bar{q}\gamma^\mu q \bar{\chi}\gamma_\mu\chi + \frac{1}{\Lambda^2} \bar{q}\gamma^\mu\gamma^5 q \bar{\chi}\gamma_\mu\chi + \frac{1}{\Lambda^2} \bar{q}\gamma^\mu q \bar{\chi}\gamma_\mu\gamma^5\chi \\ & + \frac{1}{\Lambda^2} \bar{q}\gamma^\mu\gamma^5 q \bar{\chi}\gamma_\mu\gamma^5\chi \\ & + \frac{1}{\Lambda^2} \bar{q}\sigma^{\mu\nu} q \bar{\chi}\sigma_{\mu\nu}\chi \end{aligned}$$

Instead of quarks, we could have also used leptons.

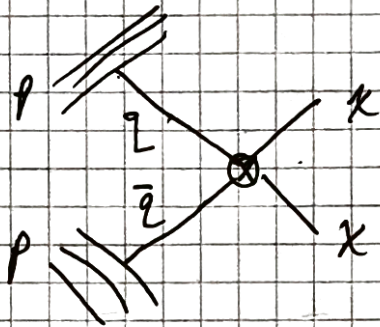
Note that scalar operators, such as $\bar{q}q + \bar{q}\gamma^5 q$, are necessarily $SU(2)$ breaking $\Rightarrow SU(2) \times U(1)_Y$ invariant operator is dim-7:

$$\frac{1}{\Lambda^3} \bar{Q} \hat{H} u_L \bar{\chi}\chi \quad \text{or} \quad \frac{1}{\Lambda^3} \bar{Q} N d_L \bar{\chi}\chi$$

Could also use Higgs fields or $1/\chi$, but have to worry about gauge invariance (≠ "minimal coupling").

(4)

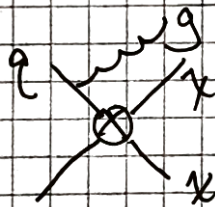
In practice, these operators lead to collider phenomenology of dark matter production.



$$\sigma \propto |M|^2 \propto \frac{1}{\Lambda^4} \cdot |C_1|^2$$

Measurements constrain strength of Wilson coefficient C_1 .
In practice, cannot "see" nothing $\Rightarrow$ need recoil object in final state to detect escaping X particles.

Collider signature is monojet + $\cancel{E}_T$, monophoton + $\cancel{E}_T$, etc., where initial state radiation gives gluon, photon, $W^\pm$, Z , etc.



ISR jet + $\cancel{E}_T$.

Main challenges: ① SM Z +jet background, with $Z \rightarrow \nu\bar{\nu}$
② EFT interpretation fails if band is too weak.