

Last time: EFTs as top-down + bottom-up tools
SM EFT

Today: EFTs in everyday use examples

① DM EFT

② Neutrino EFT + seesaw mechanism

No HW 6, but open office hours.

Bring questions or other topics to discuss.

Recall, EFTs are defined by:

① A set of low-energy/dynamical fields

② Charges/ reps. under gauge + global symmetries

③ Truncation order in power counting of UV scale cutoff Λ .

Ex. $\mathcal{O}(\frac{1}{\Lambda^2}) \rightarrow$ dim.-6 operators of SM fields,
known as SM EFT.

Distinguish top-down vs. bottom-up.

Ⓐ Start with UV completion, write corresponding EFT

Integrate of heavy dofs., truncate heavy mass scale. EX. Fermi theory

Ⓑ Start with known dofs, assign symmetries (gauge + global) to the dofs., write down all operators subject to the truncation. EX. SM EFT.

New topic: Dark Matter

We necessarily need a new field, separate from SM field content, which obeys the following 5 restrictions.

① Singlet under EM.

No electric charge.

Reason: EM interactions with CMB.
(Caveat: microcharges.)

* Small σ_{rel}
with visible
matter

② Singlet under color.

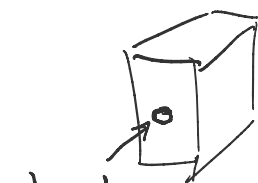
[Particle colliders would readily produce it
 $\Rightarrow$ but they have kinematic limits.]

Expect DM is color-bound state.

Reason: We expect solar system to be in a halo of DM.

Ground-based or underground direct detection.

experiment would see nuclear recoils from colored DM.



target
nucleus

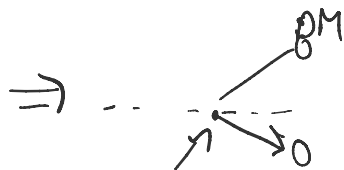
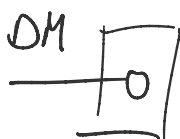
(H_2O @ Super-K

vs. LXe @ Xenon IT)

waits for DM kick.

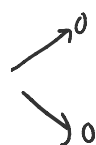
Expect dark matter density $\sim 0.3 \text{ GeV}/\text{cm}^3$

Follows Maxwell-Boltzmann distrib. $\langle v \rangle \sim 220 \text{ km/s}$.



this is
DM scattering
cross section
target

$0 \rightarrow 0 \Rightarrow$

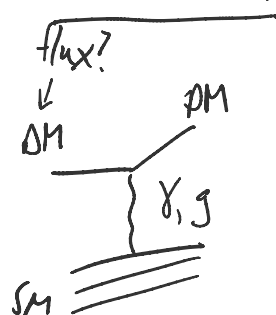


elastic scattering: mediator force.

Known force carrier: γ & $\pi^{0/2}$
 $\hookrightarrow$ color

Or NP mediator.

microscopically:



v -background for direct detection

|

...

"

.

DM

...

...

ν -background for direct detection SM —

Not currently an issue: $\sigma_{\text{nuclear recoil}}^{\text{DM}} \gtrsim 10^3 \frac{\sigma_{\nu\text{-recoil}}}{10^{-50} \text{ cm}^2}$

→ In principle, ν -flux is known. Two sources $\Rightarrow$ solar neutrinos. 8 MeV. MSW ν .
Flux: # ν per s. + atmospheric ν 's from cosmic rays.
energy distrib. of ν . extra galactic ν 's: not known
Can model the expected recoil distrib. from ν 's + subtract.

③. Cosmologically stable: particle does not decay or lifetime much longer than τ_{univ} .

④. Overall relic abundance:
Today $\Omega h^2 \sim 0.12$.

Need to know ΛCDM . (See notes)

⑤. Mass range: $10^{-20} \text{ eV} \lesssim m_{\text{DM}} \lesssim 100 M_{\odot} = 10^{59} \text{ GeV}$
fuzzy DM DM halo size
Thermal DM: is $\sim \text{MeV} \lesssim m_{\text{DM}} \lesssim 100 \text{ TeV}$.

⑥. Must act gravitationally. too warm perturbativity

Given these "rules," we can suppose

DM is scalar or fermion (or vector.)

Assume DM is $SU(2)$ singlet.

So, for concreteness, write EFT for DM assuming

DM is fermionic SM gauge singlet: χ . (Majorana

DM is fermionic SM gauge singlet: χ (Majorana or Dirac)

$$\mathcal{L}_{\text{tot}} = \mathcal{L}_{\text{SM}} + \mathcal{L}_{\text{DM}} + \mathcal{L}_{\text{DM-SM EFT}}$$

Dirac $\mathcal{L}_{\text{DM}} = \bar{\chi} i \not{\partial} \chi - m \bar{\chi} \chi$

$$\mathcal{L}_{\text{DM-SM EFT}} = \left(\begin{array}{c} \text{Typically Lorentz invariants} \\ \text{SM gauge-invariant} \end{array} \right) \otimes \left(\begin{array}{c} \text{DM gauge-invariant} \end{array} \right) \text{ at } \mathcal{O}\left(\frac{1}{\Lambda^2}\right)$$

For simplicity: work with broken phase SM.

$$SU(3)_c \times U(1)_{\text{em}}$$

To ensure DM stability, assign $\mathbb{Z}_2$ for DM $\sim +1$, SM ~ 0 .

Flesh out possible DM structures: $\left[\begin{array}{l} (\bar{e} e)(\bar{\chi} \chi) \\ h(\bar{L} e) \tilde{h}(\bar{L} \chi) \end{array} \right]$

5 possibilities: $\bar{\chi} \chi$, $\bar{\chi} \gamma^5 \chi$, $\bar{\chi} \gamma^\mu \chi$, $\bar{\chi} \sigma^{\mu\nu} \chi$, $\bar{\chi} \gamma^\mu \gamma^5 \chi$

Complete EFT:

SM fermions w/ appropriate Lorentz covariance.
($\mathbb{Z}_2$ SM neutral)

$$[\bar{q} q, \bar{q} \gamma^5 q, \bar{q} \gamma^\mu q, \bar{q} \gamma^\mu \gamma^5 q, \bar{q} \sigma^{\mu\nu} q]$$

$$\mathcal{L}_{\text{SM-DM EFT}} = \frac{1}{\Lambda^2} \bar{q} q \bar{\chi} \chi + \frac{1}{\Lambda^2} (\bar{q} \gamma^5 q \bar{\chi} \chi + \frac{1}{\Lambda^2} \bar{q} \gamma^5 q \bar{\chi} \gamma^5 \chi + \text{swap})$$

finite set $+ \frac{1}{\Lambda^2} \bar{q} \gamma^\mu q \bar{\chi} \gamma_\mu \chi + \frac{1}{\Lambda^2} \bar{q} \gamma^\mu \gamma^5 q \bar{\chi} \gamma_\mu \chi + \dots$

finite set

$$+ \frac{1}{\Lambda^2} \bar{q} \gamma^\mu q \bar{\chi} \gamma_\mu \chi + \frac{1}{\Lambda^2} \bar{q} \gamma^\mu \gamma^5 q \bar{\chi} \gamma_\mu \chi + \dots$$

$$+ \frac{1}{\Lambda^2} \bar{q} \sigma^{\mu\nu} q \bar{\chi} \sigma_{\mu\nu} \chi$$

Could also write leptonic bilinears.

What about other SM fields?

In broken phase, have $G_{\mu\nu}^a$, h , $A_\mu / F_{\mu\nu} + W_\mu^{+-} + Z_\mu$.

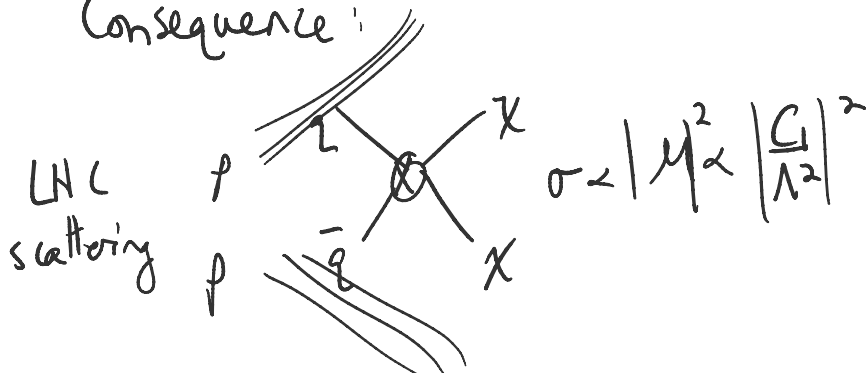
$G_{\mu\nu}^a$ does not give anything at dim-6.
 $\hookrightarrow$ not gauge inv.

Similar to $W^{+-} + Z$, since need SU(2) invariance in UV.

$$\frac{1}{\Lambda^4} \bar{q} \gamma^\mu G_{\mu\nu}^{(a)} q \bar{\chi} \sigma_{\mu\nu} \chi$$

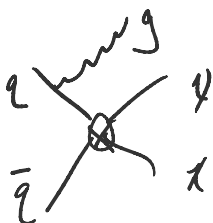
Can proceed to use EFT op. as effective int. at experiments to predict new phenomena & constrain/discover interactions.

Consequence:



$\sigma(pp \rightarrow \text{nothing}) \rightarrow$ not useful

Instead $\Rightarrow$



$\sigma(pp \rightarrow \text{jet} + \cancel{E_T}) = \text{"monojet" search}$

$$\sigma(pp \rightarrow \text{jet} + \cancel{E_T}) = \text{"monojet" search}$$

Can also do ISR photon $\Rightarrow$ "monophoton" search

or W/Z etc. $\Rightarrow$ "mono-X" search.

But exp. limits are too poor $\Rightarrow$ need simplified models.