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Theoretische Elementarteilchenphysik

Quantum Field Theory II.

Lecture 1.

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Welcome + Announcements

TAs: Julien Laux + Alexey Kivel

Exams: Oral, 50% of HW points required to register.

reader.uni-mainz.de for HW, syllabus, solutions, notes.

Introduction + Motivation to QFT II.

Physics - central goal is making quantitative predictions about natural phenomena.

Earliest example: parabolic motion in classical mechanics.

Eqn. from Newton's laws allowed prediction of speed + position of thrown ball.

Quantum physics: Pushing to apply classical knowledge to new regime of atomic systems, for example, led to failure of ability to measure $\hat{x}$ and $\hat{p}$ simultaneously
 $\Rightarrow$ Heisenberg uncertainty principle $[\hat{x}, \hat{p}] = i\hbar \neq 0$
led to birth of quantum operators tied to linear algebra $\Rightarrow$ behavior of quantum operators on wavefns exactly same as linear algebra systems. Eigenvalues become quantum numbers when operator commutes w/ Hamiltonian.

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In QFT, the analogous situation occurs when moving from free quantum field theory to interacting field theory. In free QFT, we can construct Green's fns. of the appropriate Klein-Gordon and Dirac operators to solve the Hamiltonian and build arbitrary, complicated systems. The free Hamiltonian then determines the time evolution. This was explored in QFT I from the path integral approach, for example:

When we add interactions to free theory, we can typically/generically ~~cannot~~ no longer solve the Hamiltonian. As a consequence, we appeal to a procedure of perturbative field theory. Even accepting the approximations that the ground states of the free theory and the interacting theory are equivalent, we typically resum/approximate the effects of the interaction as a perturbative expansion:

Consider QED:

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \bar{\Psi}(i\not{D} - m)\Psi$$

$$\not{D} = \gamma^\mu (\partial_\mu + iQe A_\mu) \text{ covariant derivative}$$

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$$

$$Q = \text{electric charge of } \Psi$$

$$e = \text{gauge coupling} \quad \left(\alpha = \frac{e^2}{4\pi} \text{ hyperfine structure const.} \right)$$

The non-trivial interaction between Ψ and A_μ gives the following expansion for the two-pt. correlation fn. of photons:

$$\Pi^{\mu\nu}(q^2) = \text{diagram 1} + \text{diagram 2} + \dots$$

Recall $\Pi_{1\text{PI}}^{\mu\nu}$ is the bubble of Ψ with truncated photons.

③

Analyzing the 1-loop diagram, we get the integral over virtual momentum:

$$\Pi_{\mu\nu}^{(1)} \sim (-ie)^2 \int \frac{d^4k}{(2\pi)^4} \text{Tr} \left[\frac{i(\not{k} + m)}{k^2 - m^2} \gamma_\mu \frac{i(\not{k} + \not{p} + m)}{(k+p)^2 - m^2} \gamma_\nu \right] \times (-1)$$

The main issue is that integral is quadratically divergent for $k \rightarrow \infty$!

Such UV divergent integrals can be rendered finite by extracting the essential part of the UV divergence (via dimensional regularization, Pauli-Villars regularization, or a different regularization prescription), and then interpreting finite part as a physical property of the amplitude (i.e. something that can be measured experimentally.)

Caution: Mathematically, this is very dangerous. In Pauli-Villars, for example, you are essentially subtracting two infinities to get a finite answer. But you ~~are~~ should in mind that this requires a principle value prescription without which there is no control over the value of the finite answer.)

The key point about the infinities / UV divergences is that different loop calculations in a given theory can match the infinities in such a way that all of the remaining finite parts are physical properties / observable. As a specific case, in QED, the UV divergences in the photon 2pt. correlation fn., the electron 2pt. correlation fn., and the $e^-e^+\gamma$ vertex are all matched to WFR constants $Z_{1,2}$ and a vertex counterterm Z_3 .

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[The electron mass is also renormalized.]

Because the UV divergences are "universal" for observers at any energy scale (i.e. experiments running at different energies), the finite answers measured by different observers are hence related by renormalization flow. Establishing the renormalization group equation (Callen-Symanzik equation) is the first main result of this course.

We begin by summarizing the main results of fields in QFT I to fix our conventions. (Ref. Peskin & Schroeder)

Scalar fields (solns. to Klein-Gordon)

$$[\text{real}] \quad \phi(\vec{x}) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E}} (a_p + a_{-p}^\dagger) e^{-ip \cdot x}$$

$$\pi(\vec{x}) = \int \frac{d^3p}{(2\pi)^3} (-i) \sqrt{\frac{E}{2}} (a_p - a_{-p}^\dagger) e^{-ip \cdot x}$$

$$\text{with } [a_p, a_{p'}^\dagger] = (2\pi)^3 \delta^{(3)}(\vec{p} - \vec{p}')$$

$$\left. \begin{aligned} [\phi(\vec{x}), \pi(\vec{y})] &= i \delta^{(3)}(\vec{x} - \vec{y}) \\ [\phi(\vec{x}), \phi(\vec{y})] &= [\pi(\vec{x}), \pi(\vec{y})] = 0 \end{aligned} \right\} \begin{array}{l} \text{quantization} \\ \text{condition} \\ \text{for } \phi \text{ and } \pi \end{array}$$

a_p : annihilation of momentum eigenstate $\vec{p}$

$a_p^\dagger$: creation " " " "

$\phi(\vec{x})$ creates particle at position $\vec{x}$.

$$\begin{aligned} \text{Propagator: } D(x-y) &= \langle 0 | T \phi(x) \phi(y) | 0 \rangle \\ &= \int \frac{d^4p}{(2\pi)^4} \frac{i}{p^2 - m^2 + i\epsilon} e^{-ip \cdot (x-y)} \end{aligned}$$

D is Green's fun. of Klein-Gordon operator

$$(\partial^2 + m^2) \phi = -i \delta^{(4)}(x-y)$$

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Fermion fields: (Dirac eqn.)

$$\Psi(x) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} \sum_s (a_p^s u^s(p) e^{-ip \cdot x} + b_p^{s\dagger} v^s(p) e^{ip \cdot x})$$

$$\bar{\Psi}(x) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} \sum_s (b_p^s \bar{v}^s(p) e^{-ip \cdot x} + a_p^s \bar{u}^s(p) e^{ip \cdot x})$$

$$\{a_p^s, a_q^{s\dagger}\} = \{b_p^s, b_q^{s\dagger}\} = (2\pi)^3 \delta^{(3)}(\vec{p} - \vec{q}) \delta^{rs}$$

$$\{\Psi_a(\vec{x}), \Psi_b^\dagger(\vec{y})\} = \delta^{(3)}(\vec{x} - \vec{y}) \delta_{ab}$$

$$\{\Psi_a(\vec{x}), \Psi_b(\vec{y})\} = \{\Psi_a^\dagger(\vec{x}), \Psi_b^\dagger(\vec{y})\} = 0$$

$a_p^{s\dagger}$ creates fermions, $b_p^{s\dagger}$ creates antifermions.

$$s=1,2: \quad u^s(p) = \begin{pmatrix} \sqrt{p \cdot \sigma} \xi^s \\ \sqrt{p \cdot \bar{\sigma}} \xi^s \end{pmatrix} \quad v^s(p) = \begin{pmatrix} \sqrt{p \cdot \sigma} \eta^s \\ -\sqrt{p \cdot \bar{\sigma}} \eta^s \end{pmatrix}$$

$$S(x-y) = \int \frac{d^4p}{(2\pi)^4} \frac{i(\not{p} + m)}{p^2 - m^2 + i\epsilon} e^{-ip \cdot (x-y)}$$

$$= \langle 0 | T \Psi(x) \bar{\Psi}(y) | 0 \rangle$$

Again, $S(x-y)$ is Green's fun. for Dirac operator.

Cross sections

$$d\sigma = \frac{1}{2E_A 2E_B} \frac{1}{|v_A - v_B|} \left(\prod_f \frac{d^3p_f}{(2\pi)^3} \frac{1}{2E_f} \right) |\mathcal{M}(p_A, p_B \rightarrow \{p_f\})|^2$$

$$\cdot (2\pi)^4 \delta^{(4)}(p_A + p_B - \sum_f p_f)$$

Width

$$d\Gamma = \frac{1}{2m_A} \left(\prod_f \frac{d^3p_f}{(2\pi)^3} \frac{1}{2E_f} \right) |\mathcal{M}(m_A \rightarrow \{p_f\})|^2$$

$$\cdot (2\pi)^4 \delta^{(4)}(p_A - \sum_f p_f)$$

S-matrix. Takes in states to out states in Hilbert ~~state~~ space.

Necessarily unitary. $S = \mathbb{1} + iT$

$$\langle \vec{p}_1, \vec{p}_2, \dots | iT | \vec{k}_A, \vec{k}_B \rangle = (2\pi)^4 \delta^{(4)}(k_A + k_B - \sum p_f) i \mathcal{M}(k_A, k_B \rightarrow \{p_f\})$$

In calculations, Feynman diagrams represent $\mathcal{M}$.

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Requires interactions

Mass terms, e.g. $\frac{m^2}{2} \phi^2$, $m \bar{\psi} \psi$, are simplistic interactions

In $D=4$, three or more fields gives non-trivial interactions.

Example: $g \bar{e} \gamma_\mu A^\mu e$ from QED.

$y H \bar{L}_L e + h.c.$ from Standard Model

$\frac{\lambda}{4!} \phi^4$ from scalar field theory.

Begin renormalization.

Critical issue: In an interacting theory, how do you relate the eigenstates of the interacting Hamiltonian with eigenstates of free Hamiltonian?

Recall $\langle 0 | T \phi(x) \phi(y) | 0 \rangle$: 2 pt. correlation fn. in free theory.
(propagator) - amplitude for particle to go from y to x .
 $|0\rangle$: ground state of free
 $|\Omega\rangle$: ground state of interacting
 $\langle \Omega | T \phi(x) \phi(y) | \Omega \rangle$: 2 pt. correlation fn. in interacting theory.

Schwartz
p. 70-74

Generalize to n -particles: defined for asymptotic free states.

$\langle f | M | i \rangle$ is S -matrix element with δ -fn. taken out.

$$\langle f | S - 1 | i \rangle = i \int_{t=-\infty}^{t=+\infty} (2\pi)^4 \delta^{(4)}(\Sigma p) \mathcal{M}$$

$$|i\rangle = \sqrt{2E_1} \sqrt{2E_2} a_{p_1}^\dagger(-\infty) a_{p_2}^\dagger(-\infty) |\Omega\rangle$$

$$|f\rangle = \sqrt{2E_3} \dots \sqrt{2E_n} a_{p_3}^\dagger(+\infty) a_{p_n}^\dagger(+\infty) |\Omega\rangle$$

Assume $|f\rangle \neq |i\rangle$, so discard 1.

$$\langle f | S | i \rangle = 2^{n/2} \sqrt{E_1 \dots E_n} \langle \Omega | a_{p_3}(\infty) \dots a_{p_n}(\infty) a_{p_1}^\dagger(-\infty) a_{p_2}^\dagger(-\infty) | \Omega \rangle$$

To progress, need to prove

$$i \int d^4x e^{ip \cdot x} (\square + m^2) \phi(x) = \sqrt{2E_p} (a_p(\infty) - a_p(-\infty))$$

for asymptotic free states. Assume fields die off at

$|\vec{x}| = \pm\infty$, so can integrate by parts in $\vec{x}$.

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$$i \int d^4x e^{ip \cdot x} (\square + m^2) \phi(x) = i \int d^4x e^{ip \cdot x} (\partial_t^2 - \vec{\partial}_x^2 + m^2) \phi(x)$$

$$\begin{aligned} \text{IBP, require } \left(x \phi e^{-ip \cdot x} \right) \Big|_{-\infty}^{\infty} = 0 &= i \int d^4x e^{ip \cdot x} (\partial_t^2 + \vec{p}^2 + m^2) \phi(x) \\ &= i \int d^4x e^{ip \cdot x} (\partial_t^2 + E_p^2) \phi(x) \end{aligned}$$

$$\begin{aligned} \text{Now use } \partial_t [e^{ip \cdot x} (i\partial_t + E_p) \phi(x)] &= [iE_p e^{ip \cdot x} (i\partial_t + E_p) + e^{ip \cdot x} (i\partial_t^2 + E_p \partial_t)] \phi \\ &= i e^{ip \cdot x} (\partial_t^2 + E_p^2) \phi(x) \end{aligned}$$

independent of B.C.s.

$$\begin{aligned} \text{Together, } i \int d^4x e^{ip \cdot x} (\square + m^2) \phi(x) &= \int d^4x \partial_t [e^{ip \cdot x} (i\partial_t + E_p) \phi(x)] \\ &= \int dt \partial_t [e^{iE_p t} \int d^3x e^{-i\vec{p} \cdot \vec{x}} (i\partial_t + E_p) \phi(x)] \end{aligned}$$

Integrand is total derivative, evaluate at bdy $t = \pm \infty$.

Note $a_p(t) + a_p^\dagger(t)$ operators are time-independent at $t \rightarrow \pm \infty$

Calculate

$$\begin{aligned} &\int d^3x e^{-i\vec{p} \cdot \vec{x}} (i\partial_t + E_p) \phi(x) \\ &= \int d^3x e^{-i\vec{p} \cdot \vec{x}} (i\partial_t + E_p) \int \frac{d^3k}{(2\pi)^3} \frac{1}{\sqrt{2E_k}} (a_k(t) e^{-ik \cdot x} + a_k^\dagger(t) e^{ik \cdot x}) \\ \star &= \int \frac{d^3k}{(2\pi)^3} \int d^3x \left[\frac{(E_k + E_p)}{\sqrt{2E_k}} a_k(t) e^{-ik \cdot x} e^{-i\vec{p} \cdot \vec{x}} + \frac{(-E_k + E_p)}{\sqrt{2E_k}} a_k^\dagger(t) e^{ik \cdot x} e^{-i\vec{p} \cdot \vec{x}} \right] \end{aligned}$$

using $\partial_t a_k(t) = 0$ for bdy's.

$$\text{Now, } \int d^3x e^{-ik \cdot x} e^{-i\vec{p} \cdot \vec{x}} = e^{-iE_k t} \delta^{(3)}(\vec{p} - \vec{k}) (2\pi)^3$$

and similar for second term:

$$\int d^3x e^{+ik \cdot x} e^{-i\vec{p} \cdot \vec{x}} = e^{iE_k t} \delta^{(3)}(\vec{p} + \vec{k}) (2\pi)^3$$

$$\text{We get } \star = \sqrt{2E_p} a_p(t) e^{-iE_p t}$$

$$\begin{aligned} \text{So, } i \int d^4x e^{ip \cdot x} (\square + m^2) \phi(x) &= \int dt \partial_t [\dots] \\ &= \sqrt{2E_p} (a_p(\infty) - a_p(-\infty)) \end{aligned}$$

Also, by h.c.

$$-i \int d^4x e^{-ip \cdot x} (\square + m^2) \phi(x) = \sqrt{2E_p} (a_p^\dagger(\infty) - a_p^\dagger(-\infty))$$

↑
real field

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$$\begin{aligned}
\text{So, } \langle f | S | i \rangle &= \sqrt{2^n} \sqrt{E_1 \dots E_n} \langle \Omega | a_{p_3}(\infty) \dots a_{p_n}(\infty) a_{p_1}^\dagger(-\infty) a_{p_2}^\dagger(-\infty) | \Omega \rangle \\
&= \sqrt{2^n} \sqrt{E_1 \dots E_n} \langle \Omega | T \{ a_{p_3} \dots a_{p_n} a_{p_1}^\dagger a_{p_2}^\dagger \} | \Omega \rangle \\
&= \sqrt{2^n} \sqrt{E_1 \dots E_n} \langle \Omega | T \{ [a_{p_3}(\infty) - a_{p_3}(-\infty)] \dots [a_{p_n}(\infty) - a_{p_n}(-\infty)] \\
&\quad [a_{p_1}^\dagger(-\infty) - a_{p_1}^\dagger(+\infty)] [a_{p_2}^\dagger(-\infty) - a_{p_2}^\dagger(+\infty)] \} | \Omega \rangle
\end{aligned}$$

All new a_i will shuffle to annihilate on vacuum.

$$\begin{aligned}
\langle f | S | i \rangle &= \langle p_3 \dots p_n | S | p_1 p_2 \rangle \\
&= \left[i \int d^4 x_1 e^{-i p_1 x_1} (\Box_1 + m^2) \right] \dots \left[i \int d^4 x_n e^{i p_n x_n} (\Box_n + m^2) \right] \\
&\quad \langle \Omega | T \{ \phi(x_1) \dots \phi(x_n) \} | \Omega \rangle
\end{aligned}$$

Lehmann-Symanzik-Zimmermann reduction formula.

Discussion + interpretation

For S-matrix, multiply correlation fcn. by $\Box + m^2$ +
Fourier transform.

For $\phi(k)$ as free fields, $(\Box + m^2)\phi = 0$

But one-particle states have propagators $\frac{1}{\Box + m^2} \Rightarrow \frac{1}{p^2 - m^2} \rightarrow$ will
then extract residues at pole, will
differ b/c conventions (on-shell, subtraction)

Remark: unitary theories require simple poles, so correlation
fcn. can only fall off as p^{-2} for large $\vec{p}$.

Note: for fields with spin, must also multiply by spinor
[$u(p)$] or polarization vector [$\epsilon^\mu(k)$] to project
desired spin state.